



THE THREE-BODY PROBLEM IN SHAPE SPACE

Vladimir Titov
tit@astro.spbu.ru

Chair of celestial mechanics, math-mech department, SPbSU

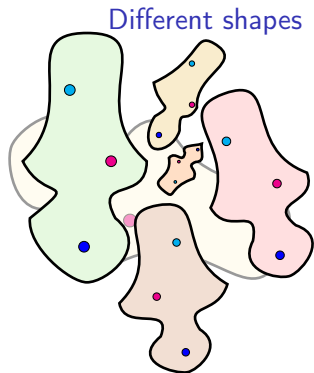
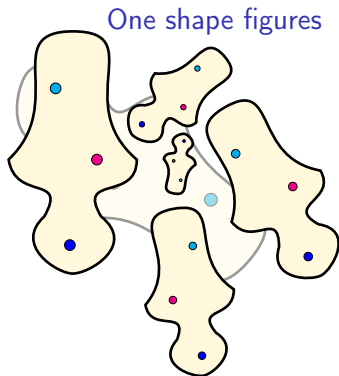
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What is it about?

- What is SHAPE (space, sphere etc)?
- Formulas and others
- Hamiltonian, energy integral and region of possible motion
- Topology of possible motion space
- Invariant configurations
- Periodic orbits
- Periodic orbits in shape space
- Regularization
- Degenerate orbits and chaos

What is SHAPE

Two planar figures have the same **shape** if one figure is exactly superposed on the other by some translation, rotation and scaling.

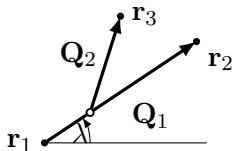


Note 1. For TBP the configuration of bodies is always triangle.

Note 2. The size is important for TBP as well.

Shape space, I. Reduction by translation ($\mathbb{R}^9 \rightarrow \mathbb{R}^6$)

$$\mathbb{R}^6 \rightarrow \mathbb{R}^4$$



$$\begin{aligned} \mathbf{Q}_1 &= \mathbf{r}_2 - \mathbf{r}_1, \\ \mathbf{Q}_2 &= \mathbf{r}_3 - \frac{m_1 \mathbf{r}_1 + m_2 \mathbf{r}_2}{m_1 + m_2}, \end{aligned}$$

$$m_1 \mathbf{x}_1 + m_2 \mathbf{x}_2 + m_3 \mathbf{x}_3 = 0.$$

In Jacobi coordinates:

$$\begin{aligned} r_{12} &= |\mathbf{Q}_1|, \\ r_{13} &= \left| \mathbf{Q}_2 + \frac{m_2}{m_1 + m_2} \mathbf{Q}_1 \right|, \\ r_{23} &= \left| \mathbf{Q}_2 - \frac{m_1}{m_1 + m_2} \mathbf{Q}_1 \right|. \end{aligned}$$

$$T = \frac{1}{2}(m_1 \dot{\mathbf{r}}_1^2 + m_2 \dot{\mathbf{r}}_2^2 + m_3 \dot{\mathbf{r}}_3^2) = \frac{1}{2} \left(\mu_1 \dot{\mathbf{Q}}_1^2 + \mu_2 \dot{\mathbf{Q}}_2^2 \right),$$

$$L = T(\dot{\mathbf{Q}}_1, \dot{\mathbf{Q}}_2) - V(\mathbf{Q}_1, \mathbf{Q}_2),$$

$$J = \mu_1 \mathbf{Q}_1 \times \dot{\mathbf{Q}}_1 + \mu_2 \mathbf{Q}_2 \times \dot{\mathbf{Q}}_2.$$

Here $\mu_1 = m_1 m_2 / (m_1 + m_2)$, $\mu_2 = m_3 (m_1 + m_2) / (m_1 + m_2 + m_3)$.

Shape space, II. Reduction by rotation

Reducton by rotation (Hopf mapping):

$$\mathbb{R}^4 \rightarrow \mathbb{R}^3 : \mathcal{S}^1 \hookrightarrow \mathcal{S}^3 \rightarrow \mathcal{S}^2 \\ (\mathcal{S}^3 \hookrightarrow \mathcal{S}^7 \rightarrow \mathcal{S}^4).$$

$$\begin{aligned} \xi_1 &= \mu_1 |\mathbf{Q}_1|^2 - \mu_2 |\mathbf{Q}_2|^2, \\ \xi_2 &= 2\sqrt{\mu_1 \mu_2} \mathbf{Q}_1 \cdot \mathbf{Q}_2, \\ \xi_3 &= 2\sqrt{\mu_1 \mu_2} \mathbf{Q}_1 \times \mathbf{Q}_2, \end{aligned}$$

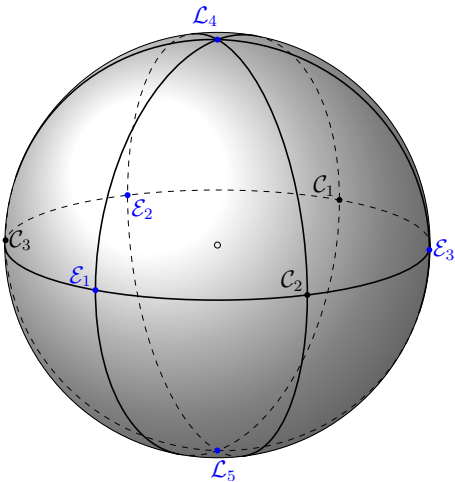
Moment of inertia:

$$\begin{aligned} I &= m_1 |\mathbf{r}_1|^2 + m_2 |\mathbf{r}_2|^2 + m_3 |\mathbf{r}_3|^2 = m_1 m_2 r_{12}^2 + m_1 m_3 r_{13}^2 + m_2 m_3 r_{23}^2 = \\ &= \mu_1 |\mathbf{Q}_1|^2 + \mu_2 |\mathbf{Q}_2|^2 = \sqrt{\xi_1^2 + \xi_2^2 + \xi_3^2}. \end{aligned}$$

Shape space, III

Shape space, $\Xi = (\xi_1, \xi_2, \xi_3)$, is the space of congruent triangles.

Shape sphere, $S^2 \subset \Xi$, is the space of similar triangles.



If $\mathbf{r}_i(t)$, $i = 1, \dots, N$, is solution of N -body problem, the following expression gives the solution as well

$$\boldsymbol{\rho}_i(t) = \lambda \mathbf{r}_i(\lambda^{-3/2}t)$$

In addition

$$\begin{aligned}\dot{\boldsymbol{\rho}}_i &= \lambda^{-1/2} \mathbf{v}_i(\lambda^{-3/2}t) \\ h' &= h/\lambda\end{aligned}$$

Expressions for mutual distances

$$r_{12}^2 = \frac{m_1 + m_2}{2m_1m_2} (\sqrt{\xi_1^2 + \xi_2^2 + \xi_3^2} + \xi_1)$$

$$r_{13}^2 = \frac{m_1 + m_3}{2m_1m_3} \sqrt{\xi_1^2 + \xi_2^2 + \xi_3^2} + \frac{m_2m_3 - m_1(m_1 + m_2 + m_3)}{2m_1m_3(m_1 + m_2)} \xi_1 + \frac{\sqrt{m_1m_2m_3(m_1 + m_2 + m_3)}}{m_1m_3(m_1 + m_2)} \xi_2$$

$$r_{23}^2 = \frac{m_2 + m_3}{2m_2m_3} \sqrt{\xi_1^2 + \xi_2^2 + \xi_3^2} + \frac{m_1m_3 - m_2(m_1 + m_2 + m_3)}{2m_2m_3(m_1 + m_2)} \xi_1 - \frac{\sqrt{m_1m_2m_3(m_1 + m_2 + m_3)}}{m_2m_3(m_1 + m_2)} \xi_2$$

$$\begin{aligned}
T &= \frac{4J^2 + \dot{\xi}_1^2 + \dot{\xi}_2^2 + \dot{\xi}_3^2}{8\sqrt{\xi_1^2 + \xi_2^2 + \xi_3^2}}, \\
J &= \mu_1 \mathbf{Q}_1 \times \dot{\mathbf{Q}}_1 + \mu_2 \mathbf{Q}_2 \times \dot{\mathbf{Q}}_2, \\
&= \sqrt{\xi_1^2 + \xi_2^2 + \xi_3^2} \frac{d\lambda}{dt} + \frac{\xi_2 \frac{d\xi_3}{dt} - \xi_3 \frac{d\xi_2}{dt}}{2(\sqrt{\xi_1^2 + \xi_2^2 + \xi_3^2} + \xi_1)}
\end{aligned}$$

Clearly, the angular momentum J is conjugated to the angle λ . If we know ξ_1 , ξ_2 , ξ_3 , we can obtain this angle λ from quadrature:

$$\lambda(t) = \int_0^t \frac{\partial R}{\partial J} d\tau = \int_0^t \frac{J - \frac{\xi_2 \dot{\xi}_3 - \xi_3 \dot{\xi}_2}{2\sqrt{\xi_1^2 + \xi_2^2 + \xi_3^2} + \xi_1}}{\sqrt{\xi_1^2 + \xi_2^2 + \xi_3^2}} d\tau.$$

The system is conservative, so we have energy integral:

$$T - U = \frac{4J^2 + \dot{\xi}_1^2 + \dot{\xi}_2^2 + \dot{\xi}_3^2}{8\sqrt{\xi_1^2 + \xi_2^2 + \xi_3^2}} - U(\xi_1, \xi_2, \xi_3) = h$$

The region of possible region

Proposition (Arnold et al, 1.7)

We have the inequality: $J^2 \leq 2IT$

$$J^2 = \left| \sum m_i (\mathbf{r}_i \times \mathbf{v}_i) \right|^2 \leq \left(\sum m_i |\mathbf{r}_i| |\mathbf{v}_i| \right)^2 \leq \left(\sum m_i r_i^2 \right) \left(\sum m_i v_i^2 \right) = I \cdot 2T.$$

Constraints (the kinetic integrals)

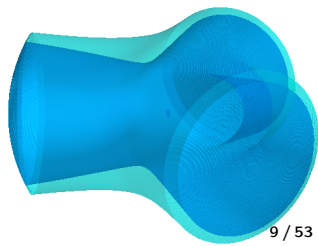
$$\Gamma = \left\{ \mathbf{r} = (\mathbf{r}_1, \dots, \mathbf{r}_n) \in \mathbb{R}^{3n} : \sum m_i \mathbf{r}_i = 0 \right\}.$$

In planar three body problem (Proposition 1.8)

$$B_{J,h} = \left\{ \mathbf{r} \in \Gamma : U + \frac{J^2}{2I} \leq h \right\},$$

in spacial one

$$B_{J,h} \subset \left\{ \mathbf{r} \in \Gamma : U + \frac{J^2}{2I} \leq h \right\} \subset \Gamma.$$



Energy integral, Sundman inequality and zero velocity surface

$$\begin{aligned} \frac{J^2}{2I} - U(\xi_1, \xi_2, \xi_3) - h &\leq \\ \frac{J^2}{2I} + \dot{I}^2/(8I) - U(\xi_1, \xi_2, \xi_3) - h &\leq \\ \frac{4J^2 + \dot{\xi}_1^2 + \dot{\xi}_2^2 + \dot{\xi}_3^2}{8\sqrt{\xi_1^2 + \xi_2^2 + \xi_3^2}} - U(\xi_1, \xi_2, \xi_3) - h &= 0, \end{aligned}$$

$$\frac{J^2}{2\sqrt{\xi_1^2 + \xi_2^2 + \xi_3^2}} - U(\xi_1, \xi_2, \xi_3) = h,$$

$$U(\xi_1, \xi_2, \xi_3) = \frac{1}{\sqrt[4]{\xi_1^2 + \xi_2^2 + \xi_3^2}} D(\varphi, \theta).$$

$m_1 = m_2 = m_3 = 1$: $O\xi_3$ и $O\xi_1$

$$J = 0: \quad U(0, 0, \xi_3) = 3/\sqrt{\xi_3} = -h = \frac{1}{2}, \quad \rightarrow \quad \xi_3 = 36$$

$$U(\xi_1, 0, 0) = 1/\sqrt{2\xi_1} + 2/\sqrt{\xi_1/2}, \quad \rightarrow \quad \xi_1 = 50.$$

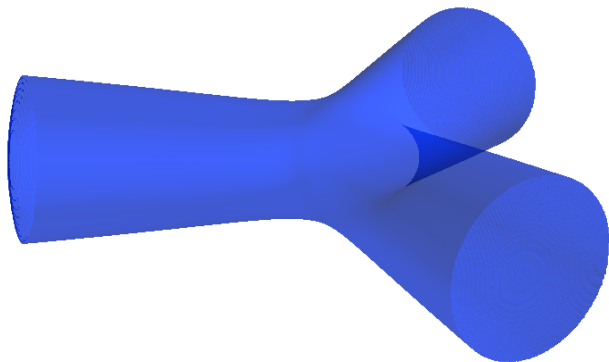
$$J \neq 0: \quad U(0, 0, \xi_3) - \frac{J^2}{2\xi_3} = 3/\sqrt{\xi_3} - \frac{J^2}{2\xi_3} = \frac{1}{2} \quad \rightarrow$$

$$\xi_3 \in \left[\left(3 - \sqrt{9 - J^2} \right)^2, \left(3 + \sqrt{9 - J^2} \right)^2 \right].$$

$$U(\xi_1, 0, 0) - \frac{J^2}{2\xi_1} = 1/\sqrt{2\xi_1} + 2/\sqrt{\xi_1/2} - \frac{J^2}{2\xi_1} = \frac{1}{2} \quad \rightarrow$$

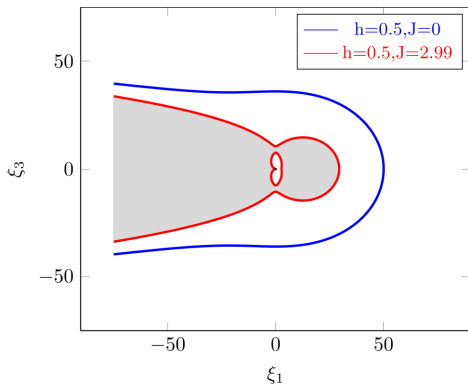
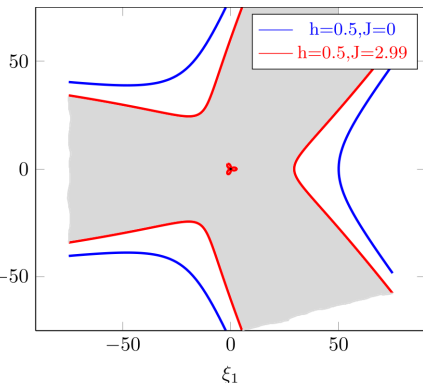
$$\xi_1 \in \left[\left(5 - \sqrt{25 - 2J^2} \right)^2 / 2, \left(5 + \sqrt{25 - 2J^2} \right)^2 / 2 \right].$$

Zero velocity surface at $J = 0$ ($h = -\frac{1}{2}$)

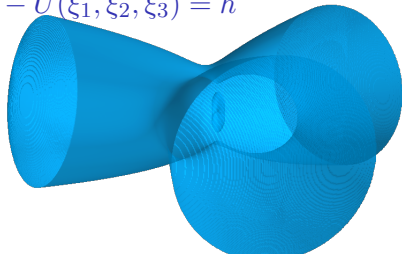


$$U(\xi_1, \xi_2, \xi_3) = -h$$

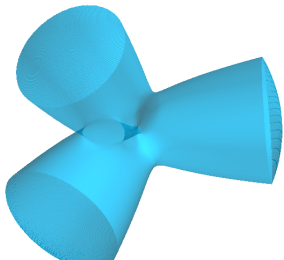
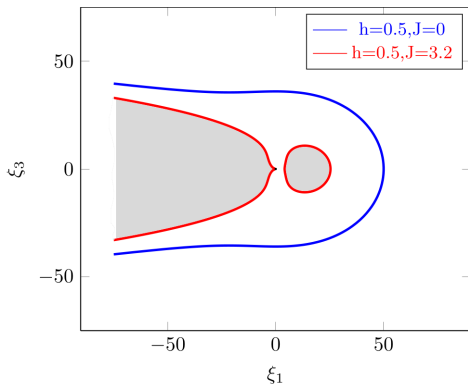
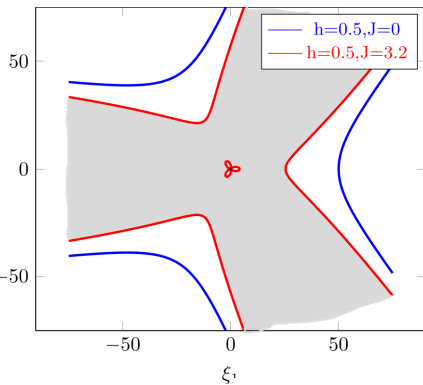
Zero velocity surface at $J = 2.99$



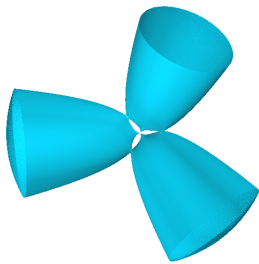
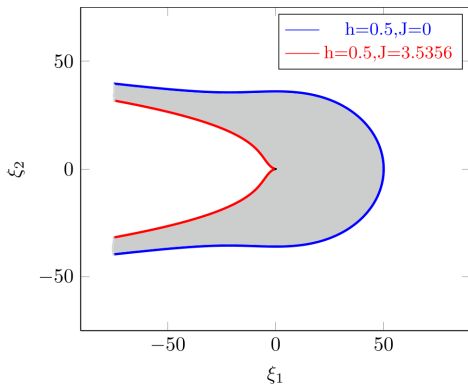
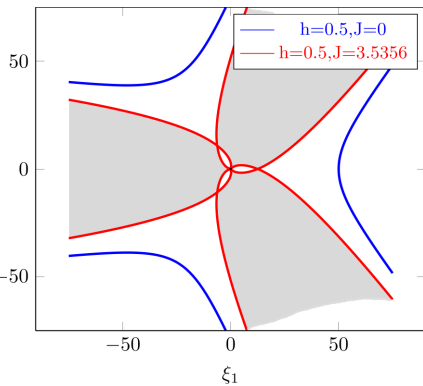
$$\frac{J^2}{2\sqrt{\xi_1^2 + \xi_2^2 + \xi_3^2}} - U(\xi_1, \xi_2, \xi_3) = h$$



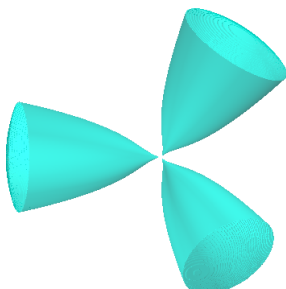
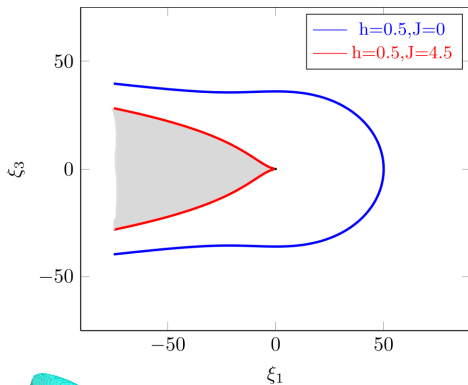
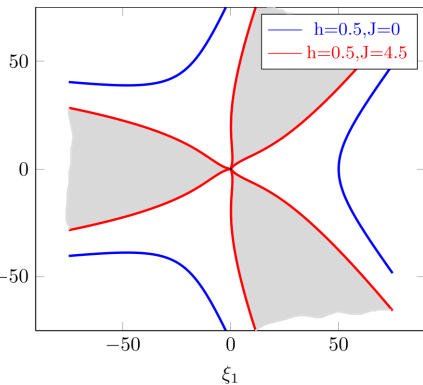
Zero velocity surface at $J = 3.2$)



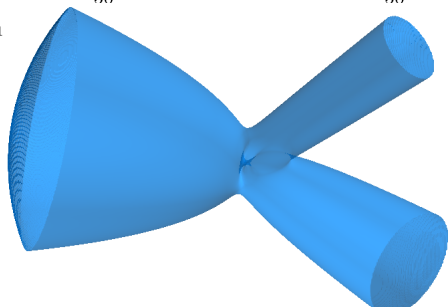
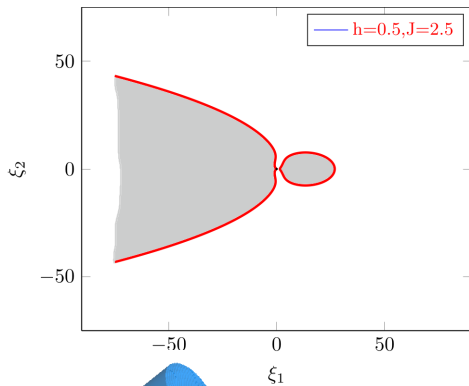
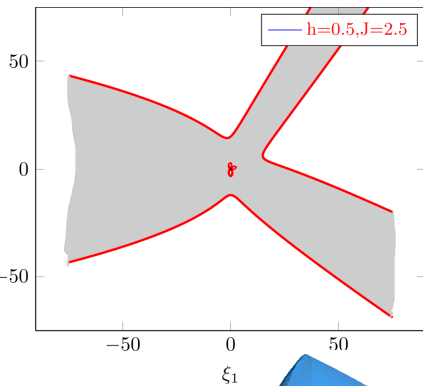
Zero velocity surface at $J = 3.53$



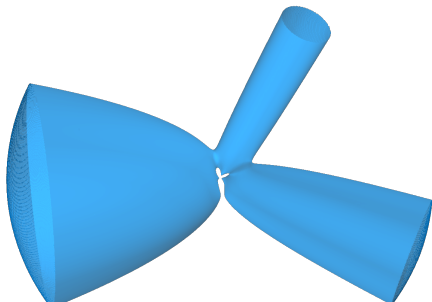
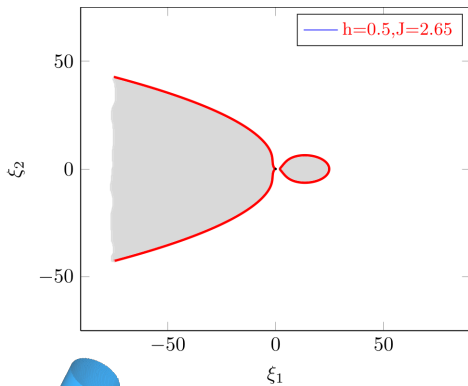
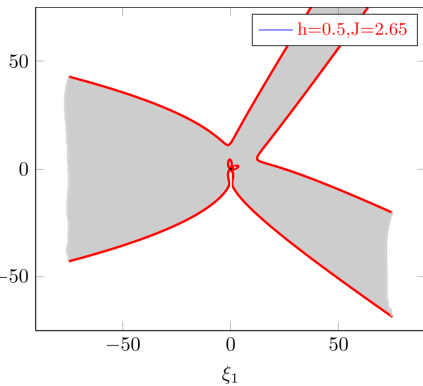
Zero velocity surface at $J = 4.5$ ($J > 5/\sqrt{2}$)



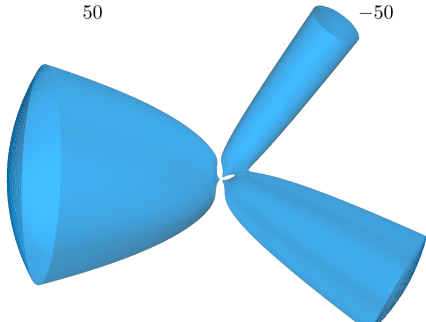
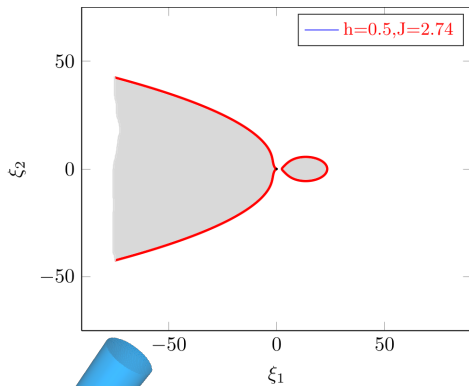
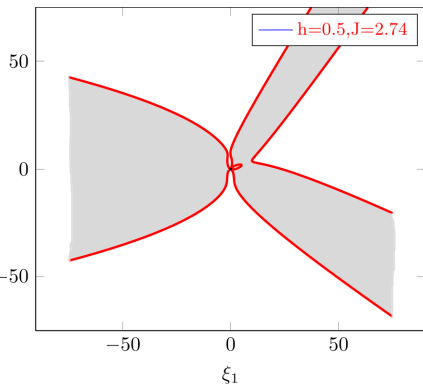
$$m_1 = 2m_2 = 4m_3 = 12/7, \quad h = -1/2, \quad J = 2.5$$



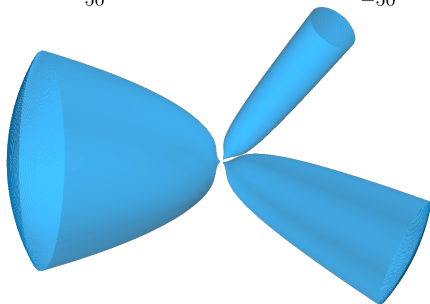
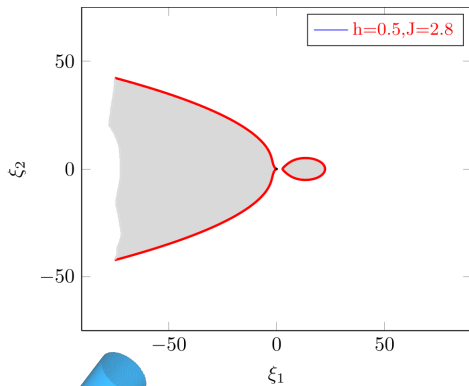
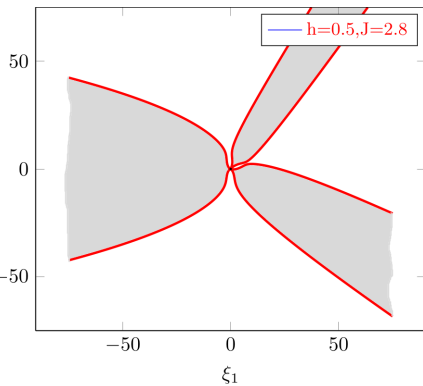
$$m_1 = 2m_2 = 4m_3 = 12/7, J = 2.65$$



$$m_1 = 2m_2 = 4m_3 = 12/7, J = 2.74$$



$$m_1 = 2m_2 = 4m_3 = 12/7, J = 2.8$$



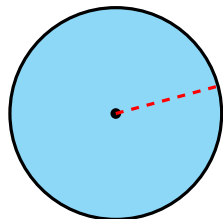
Topology of possible motion space

- 1 The available is bounded by a surface with three branches. Motion is possible inside this space with the exception of the punctured point, the origin of coordinates. As J grows the outer surface with three branches decreases, and the punctured point becomes a surface whose cross-section with the equator plane resembles a trefoil, and the cross-section with the meridian plane has the shape «frigole»,
- 2 with growth of J , the outer and inner surfaces join together and a hole forms in the outer surface,
- 3 with further growth of J , two branches are first separated, the space of possible motion is still connected,
- 4 but as J increases, one branch separates from the other two, and finally,
- 5 beginning from a certain value of J , we have three separate areas of possible movement.

Zero velocity surface (2BP)

The energy integral

$$T - V = \frac{\dot{\mathbf{r}}^2}{2} - \frac{1}{r} = h$$
$$r \leq -\frac{1}{h}, \quad h < 0.$$



Adding the angular momentum integral

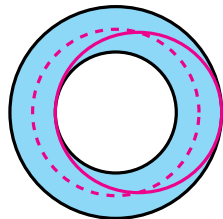
$$\mathbf{r} \times \dot{\mathbf{r}} = \mathbf{J} = r^2 \dot{\theta}$$

yields

$$\frac{2}{r} + 2h - \frac{J^2}{r^2} \geq 0$$

or

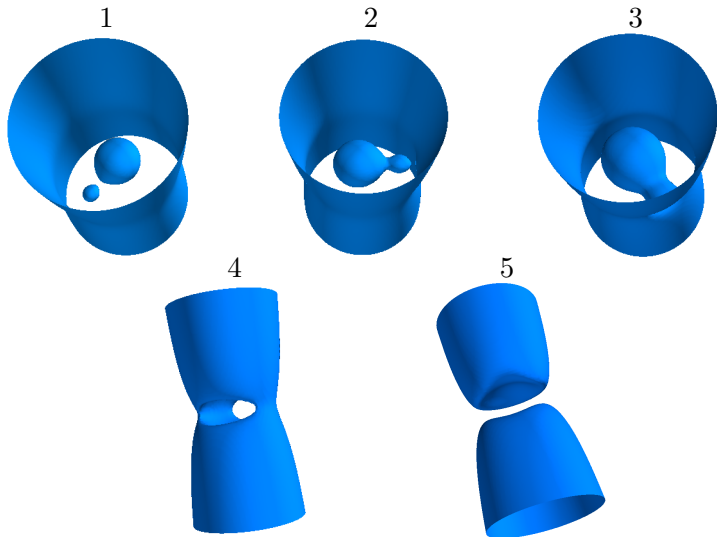
$$\begin{cases} r_{\min} \leq r \leq r_{\max} & \text{if } h < 0, \\ r > r_{\min} & \text{if } h \geq 0 \end{cases}$$



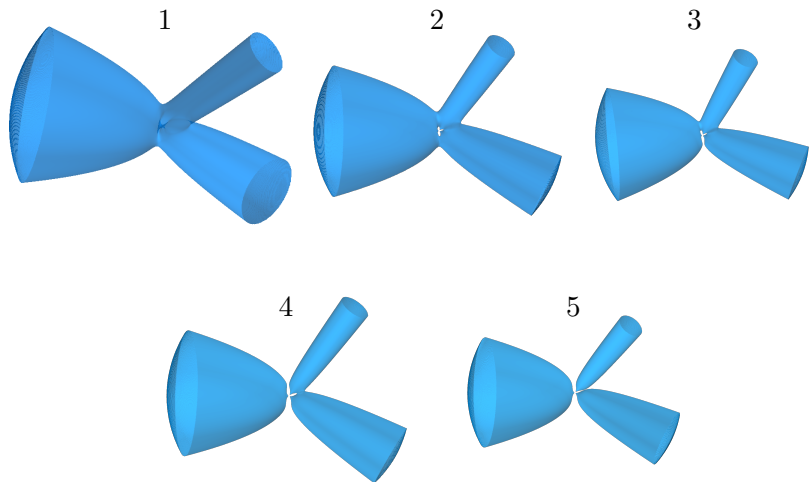
Cartesian relative or barycentric coordinate system

Zero velocity surface (Hill, CR3BP)

The Hill surface, depending on the Jacobi constant, consists of 1, 2, or 3 pieces. One or two pieces are non-compact.

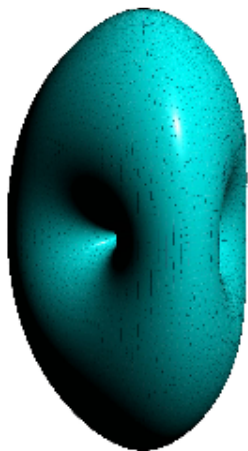


Zero velocity surface (general three-body problem)



Final motions, hierarchical systems

Inner surface



Shape space, spherical coordinates

$$\begin{aligned}\xi_1 &= \rho \cos \varphi \cos \theta, & r_{12}^2 &= \frac{m_1 + m_2}{2m_1 m_2} \rho(1 + \cos \varphi \cos \theta), \\ \xi_2 &= \rho \sin \varphi \cos \theta, & r_{13}^2 &= \frac{m_1 + m_3}{2m_1 m_3} \rho(1 - \cos(\varphi - \varphi_{13}) \cos \theta), \\ \xi_3 &= \rho \sin \theta, & r_{23}^2 &= \frac{m_2 + m_3}{2m_2 m_3} \rho(1 - \cos(\varphi - \varphi_{23}) \cos \theta).\end{aligned}$$

$$V(\rho, \theta, \varphi) = \frac{1}{\sqrt{\rho}} \left(\frac{\nu_{12}}{\sqrt{1 + \cos \theta \cos \varphi}} + \frac{\nu_{13}}{\sqrt{1 - \cos \theta \cos(\varphi - \varphi_{13})}} + \frac{\nu_{23}}{\sqrt{1 - \cos \theta \cos(\varphi - \varphi_{23})}} \right) = \frac{1}{\sqrt{\rho}} D(\theta, \varphi),$$

Invariant configurations ($\theta = \text{const}, \varphi = \text{const}$):

$$\begin{aligned}\frac{\partial D(\varphi, \theta)}{\partial \theta} &= 0, \\ \frac{\partial D(\varphi, \theta)}{\partial \varphi} &= 0\end{aligned}$$

Motion with invariant configuration

$$V = \frac{C_1}{\sqrt{\rho}}.$$

Lagrange–Jacobi identity

$$\ddot{I} = 2 \left(\frac{C_1}{\sqrt{I}} + 2h \right),$$

reduces to ($r^2 = I$, $\frac{dt}{d\tau} = r = \sqrt{I}$) κ

$$r'' = 2hr + C_1.$$

The solution is ($h < 0$)

$$r = \sqrt{I} = A \cos(n\tau - \vartheta) - \frac{C_1}{2h}, \quad n = \sqrt{-2h}.$$

Periodic orbits

Finite symmetries. For planar 3BP there are 10 finite symmetry groups only. Three of them are considered here for illustration the trajectories in shape space: the **dihedral group** (D_{12} , simple choreography), **2 – 1 choreographies** and **Line** symmetry.

Periodic orbits are searched as minimizer of action functional

$$\mathcal{A} = \int_{t_1}^{t_2} L(\mathbf{q}_i, \dot{\mathbf{q}}_i, t).$$

in the form

$$\begin{aligned}x_j(t) &= C_x^0 + \sum_{i=1} C_{x_i}^j \cos it + S_{x_i}^j \sin it \\y_j(t) &= C_y^0 + \sum_{i=1} C_{y_i}^j \cos it + S_{y_i}^j \sin it,\end{aligned}$$

Figure-Eight

$$\begin{aligned}x(t) &= 1.0959 \sin t & - & 0.0253 \sin 5t & - & 0.0058 \sin 7t \\ & + 0.0004 \sin 11t & + & 0.0001 \sin 13t \\y(t) &= 0.3373 \sin 2t & + & 0.0557 \sin 4t & - & 0.0030 \sin 8t \\ & - 0.0008 \sin 10t & + & 0.0001 \sin 14t\end{aligned}$$

Numerically found by Cr. Moore
in "Braids in classical dynamics",
Phys. Rev. Lett. 1993, **70**.

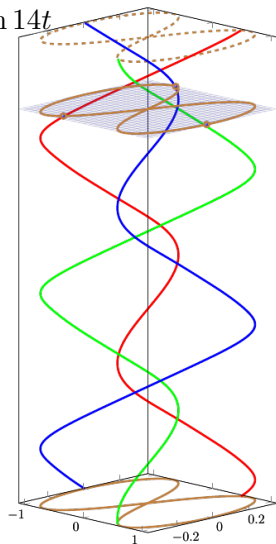
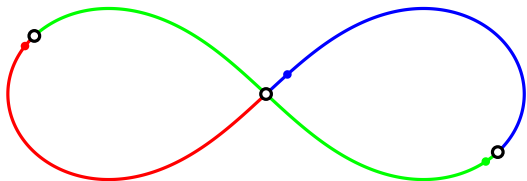
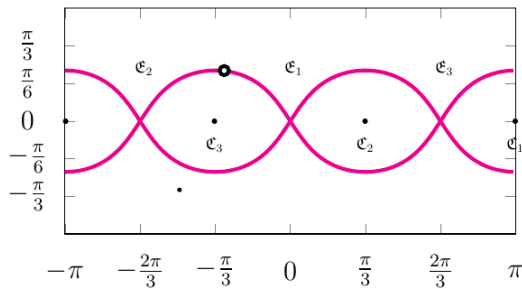
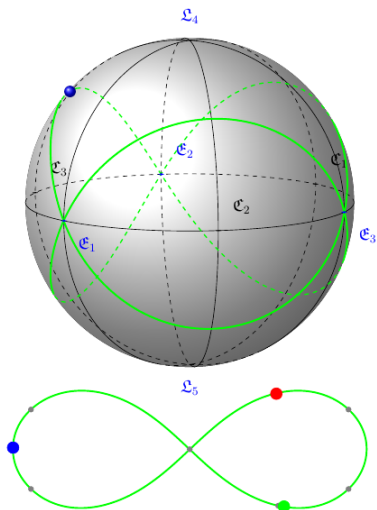
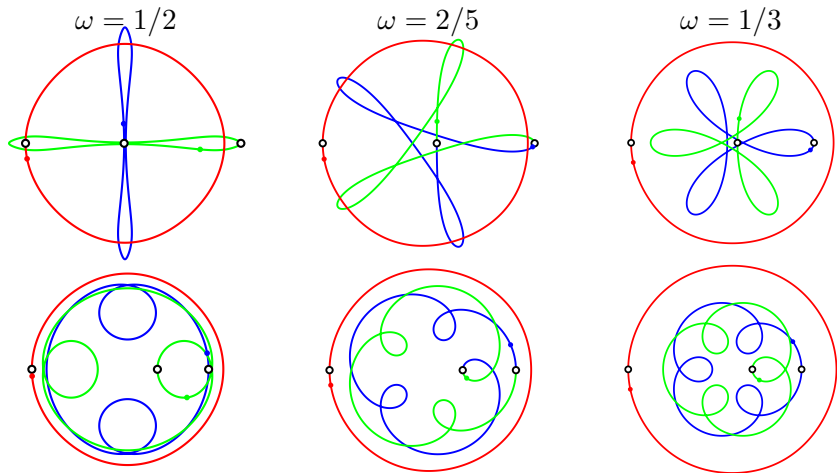


Figure-Eight on the shape sphere



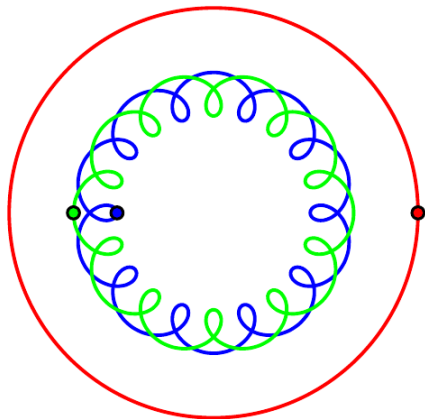
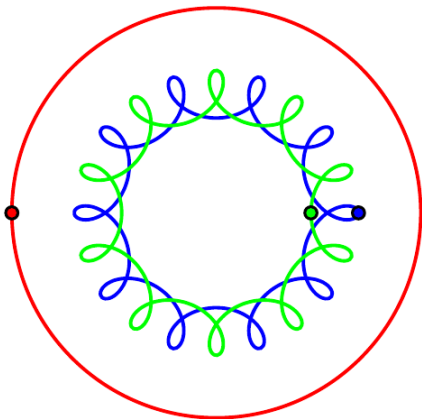
2 – 1 choreography

Two bodies of equal mass move along the same trajectory with a phase lag π . A group have a Rotating Circle Property. An orbit in an inertial system is a minimizer found in a system rotating with angular velocity ω .



Tight binaries

$$\omega = 1/3, k = 5$$

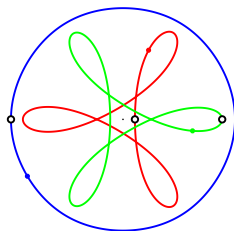
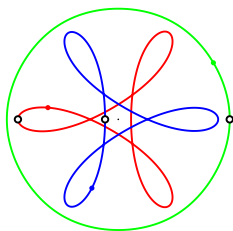
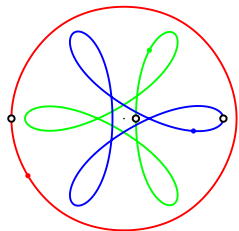
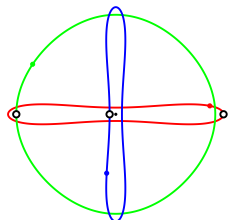
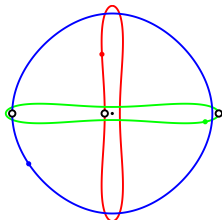
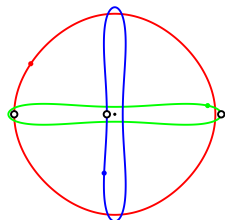


2 – 1 symmetry choreography

$m_1 = m_2 = 0.95, m_3 = 1.1$					
A	E	$ C $	ω	$[I_{\min}, I_{\max}]$	St
10.61083	-0.562922	1.73204	1/5	[13.520,18.706]	+
11.87886	-0.630193	1.34061	1/3	[7.646,7.695]	+
12.41405	-0.658586	1.22094	2/5	[6.446,6.518]	+
12.43822	-0.850687	3.17929	1/5	[13.037,13.062]	+
13.13826	-0.697007	1.09433	1/2	[5.463,5.580]	+
14.90941	-0.790968	2.76171	1/3	[6.779,6.847]	+
16.03507	-0.850687	2.61695	2/5	[5.352,5.457]	+
16.57031	-0.879082	2.44831	1/3	[3.869,3.957]	-
17.61955	-0.934746	2.43060	1/2	[3.967,4.154]	-
19.78460	-1.049610	1.57727	1/3	[6.501,6.503]	+
21.89957	-1.161810	2.58582	1/3	[6.441,6.443]	+
25.74992	-1.366082	1.65989	1/3	[6.380,6.381]	+
27.53447	-1.460752	2.51159	1/3	[6.362,6.363]	+
$m_1 = m_2 = 1.05, m_3 = 0.9$					
12.20094	-0.647280	0.98928	1/3	[7.276,7.323]	+
15.79177	-0.837779	2.68412	1/3	[6.275,6.341]	+
16.61662	-0.881539	2.33447	1/3	[3.779,3.943]	-
Figure-eight: $m_1 = m_2 = m_3 = 1.0$					
24.37193	-1.29297	0	-	[1.973,1.982]	+

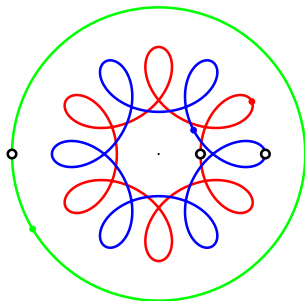
Line symmetry orbits

Three orbits with linear symmetry corresponding to cyclic permutation of masses m_1, m_2, m_3 ($\omega = 1/2$ и $\omega = 1/3$).

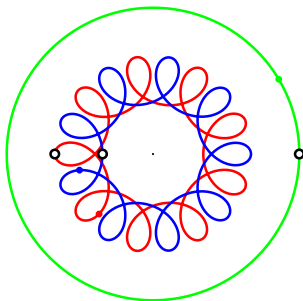


Orbits with Tight binaries

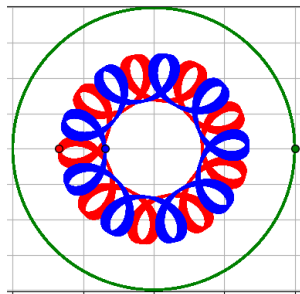
$$\omega = 1/3$$



$$k = 2$$



$$k = 3$$



Numerical integration
500 periods

Line symmetry orbits

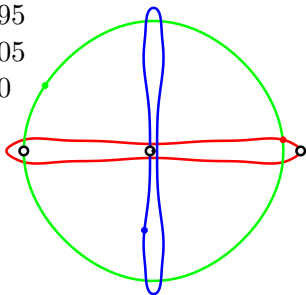
$m_1 = 0.99, m_2 = 1.01, m_3 = 1.0$					
A	E	$ J $	ω	$[I_{\min}, I_{\max}]$	St
11.42286	-0.606002	1.36301	1/4	[10.095,10.108]	+
12.04740	-0.639135	1.19429	1/3	[7.508,7.557]	+
12.06332	-0.639979	1.17690	1/3	[7.489,7.538]	+
12.07962	-0.640844	1.15915	1/3	[7.471,7.520]	+
13.15385	-0.697833	0.92132	1/2	[5.474,5.590]	+
13.15566	-0.697930	0.93926	1/2	[5.474,5.590]	+
13.15748	-0.698026	0.95484	1/2	[5.534,5.591]	+
14.06146	-0.745984	0.83708	1/3	[5.156,5.393]	-
14.08066	-0.747002	0.85327	1/3	[5.114,5.347]	-
14.09948	-0.748001	0.86909	1/3	[5.098,5.332]	-
14.55725	-0.772286	0.88706	1/4	[5.253,5.574]	-
16.64808	-0.883208	1.19288	1/3	[3.830,3.964]	-
16.76479	-0.889400	1.37020	1/3	[6.487,6.492]	+
17.80747	-0.944715	2.06327	1/4	[3.189,3.517]	-
20.59152	-1.09242	1.45497	1/3	[6.276,6.278]	+

Varying masses

$$m_1 = 0.95$$

$$m_2 = 1.05$$

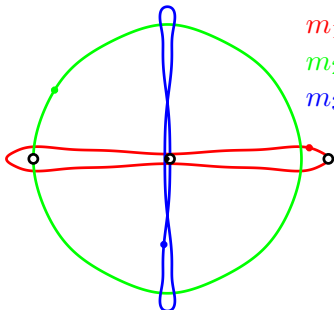
$$m_3 = 1.0$$



$$m_1 = 0.9$$

$$m_2 = 1.1$$

$$m_3 = 1.0$$

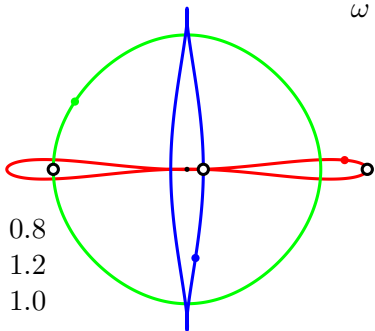


$$\omega = 1/2$$

$$m_1 = 0.8$$

$$m_2 = 1.2$$

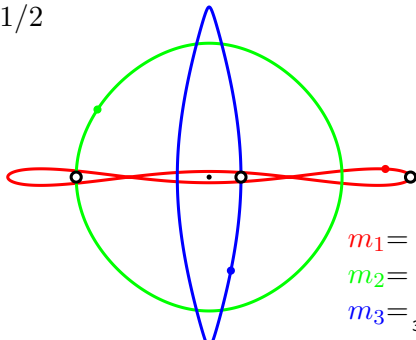
$$m_3 = 1.0$$



$$m_1 = 0.7$$

$$m_2 = 1.3$$

$$m_3 = 1.0_{37/53}$$

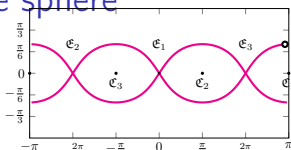


Varying masses

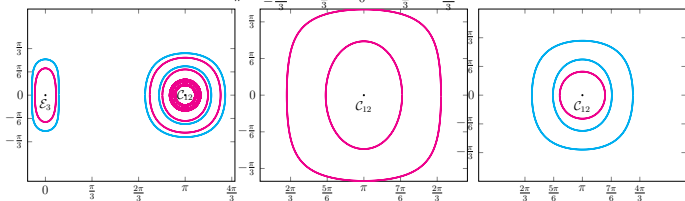
m_1	$m_1 + m_2 = 2, m_3 = 1.0, \omega = 1/2$				
	A	E	$ J $	$[I_{\min}, I_{\max}]$	St
0.99	13.15748	-0.698026	0.95484	[5.534,5.591]	+
0.95	13.15312	-0.697795	1.01964	[5.470,5.588]	+
0.9	13.12580	-0.696348	1.09648	[5.456,5.575]	+
0.8	12.99779	-0.689554	1.23654	[4.722,5.568]	+
0.7	12.77091	-0.677518	1.35872	[4.054,5.800]	+

Periodic orbits on shape sphere

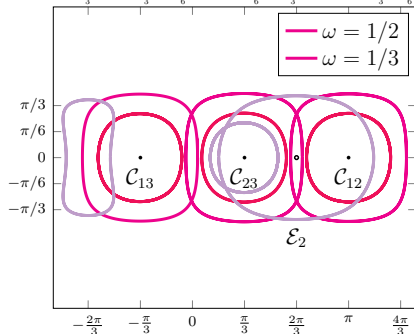
Figure-eight:



2 - 1:

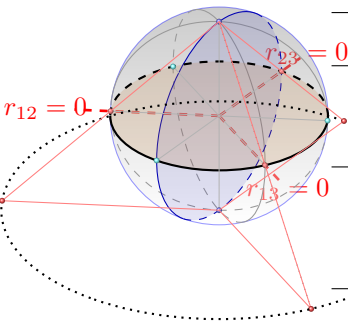


Line:



Global regularization by Lemaitre

$$\zeta = z \frac{\sqrt{8} + z^3}{1 - \sqrt{8}z^3},$$



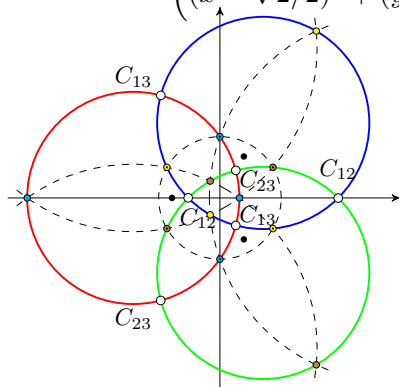
Coll.	extended complex plane z	ζ
C_{12}	$\frac{\sqrt{2}-\sqrt{6}}{2}$ $\frac{\sqrt{2}+\sqrt{6}}{2}$	-1
C_{13}	$-\frac{\sqrt{2}}{2} \begin{pmatrix} \frac{1}{2} + \frac{\sqrt{3}}{2} \\ -\frac{1}{2} + \frac{\sqrt{3}}{2} \end{pmatrix} + \frac{\sqrt{6}}{2} \begin{pmatrix} \frac{1}{2} + \frac{\sqrt{3}}{2} \\ -\frac{1}{2} + \frac{\sqrt{3}}{2} \end{pmatrix}$ $\frac{\sqrt{2}}{2} \begin{pmatrix} \frac{1}{2} + \frac{\sqrt{3}}{2} \\ -\frac{1}{2} + \frac{\sqrt{3}}{2} \end{pmatrix} - \frac{\sqrt{6}}{2} \begin{pmatrix} \frac{1}{2} + \frac{\sqrt{3}}{2} \\ -\frac{1}{2} + \frac{\sqrt{3}}{2} \end{pmatrix}$	$\frac{1}{2} - \frac{\sqrt{3}}{2} \iota$ $\frac{1}{2} + \frac{\sqrt{3}}{2} \iota$
C_{23}	$-\frac{\sqrt{2}}{2} \begin{pmatrix} \frac{1}{2} + \frac{\sqrt{3}}{2} \\ -\frac{1}{2} + \frac{\sqrt{3}}{2} \end{pmatrix} - \frac{\sqrt{6}}{2} \begin{pmatrix} \frac{1}{2} + \frac{\sqrt{3}}{2} \\ -\frac{1}{2} + \frac{\sqrt{3}}{2} \end{pmatrix}$ $\frac{\sqrt{2}}{2} \begin{pmatrix} \frac{1}{2} + \frac{\sqrt{3}}{2} \\ -\frac{1}{2} + \frac{\sqrt{3}}{2} \end{pmatrix} + \frac{\sqrt{6}}{2} \begin{pmatrix} \frac{1}{2} + \frac{\sqrt{3}}{2} \\ -\frac{1}{2} + \frac{\sqrt{3}}{2} \end{pmatrix}$	$\frac{1}{2} - \frac{\sqrt{3}}{2} \iota$ $\frac{1}{2} + \frac{\sqrt{3}}{2} \iota$

Preimages of unit circle

$$|\zeta|^2 = \zeta\bar{\zeta} = 1.$$

Direct calculations lead to the equation

$$\begin{aligned} & \left((x + \sqrt{2})^2 + y^2 - 3 \right) \\ & \left((x - \sqrt{2}/2)^2 + (y - \sqrt{6}/2)^2 - 3 \right) \\ & \left((x - \sqrt{2}/2)^2 + (y + \sqrt{6}/2)^2 - 3 \right) = 0 \end{aligned}$$



Degenerate trajectories; rectilinear orbits

$$(m_1 = m_2 = m_3 = 1)$$

$$T = \frac{\dot{\xi}_1^2 + \dot{\xi}_2^2}{8\sqrt{\xi_1^2 + \xi_2^2}}$$

Polar coordinates of (ξ_1, ξ_2)

$$\begin{aligned}\xi_1 &= \varrho^2 \cos(\varphi), \\ \xi_2 &= \varrho^2 \sin(\varphi),\end{aligned}$$

$$T = \frac{1}{2} \left(\dot{\varrho}^2 + \frac{1}{4} \varrho^2 \dot{\varphi}^2 \right).$$

$$\begin{aligned}V &= 1/r_{12} + 1/r_{13} + 1/r_{23} = \\ &= \frac{1}{\varrho} \left(\frac{1}{\sqrt{1 + \cos \varphi}} + \frac{1}{\sqrt{1 - \cos(\varphi - \pi/3)}} + \frac{1}{\sqrt{1 - \cos(\varphi - 5\pi/3)}} \right) \\ &= \frac{1}{\varrho} \frac{1 + 4 \cos \varphi}{\sqrt{1 + \cos \varphi} (2 \cos \varphi - 1)} = \frac{D(\theta)}{\rho}.\end{aligned}$$

Singular values of φ : $\varphi = \pm\pi/3, \pi$.

Rectilinear trajectories

Energy integral

$$2h = \dot{\varrho}^2 + \frac{1}{4}\varrho^2\dot{\varphi}^2 - 2V = \dot{\varrho}^2 + \frac{1}{4}\varrho^2\dot{\varphi}^2 - 2\frac{D(\varphi)}{\varrho}. \quad (1)$$

Possible motion region (in coordinates ϱ, φ):

$$\frac{D(\varphi)}{\varrho} = \frac{1}{\varrho} \frac{1 + 4 \cos \varphi}{\sqrt{1 + \cos \varphi}(2 \cos \varphi - 1)} \geq -h.$$

Zero velocity curve ($h = -1/2$) is given by

$$\varrho = \frac{2(1 + 4 \cos \theta)}{\sqrt{1 + \cos \theta}(2 \cos \theta - 1)}. \quad (2)$$

Parametrization

$$-\sqrt{2} + \sqrt{3} \cos \psi + i \sqrt{3} \sin \psi \quad \text{vs} \quad z_E = \frac{\cos E}{\sqrt{3} + \sqrt{2} \cos E} + i \frac{\sqrt{3} \sin E}{\sqrt{3} + \sqrt{2} \cos E}$$

$$\begin{aligned} \cos \varphi &= \frac{5 + \cos 4E}{7 - \cos 4E}, \\ \sin \varphi &= \frac{4\sqrt{3} \sin 2E}{7 - \cos 4E}. \end{aligned}$$

then

$$V = \frac{1}{r_{12}} + \frac{1}{r_{13}} + \frac{1}{r_{23}} = \frac{1}{\varrho} \frac{\sqrt{7 - \cos 4E} (9 + \cos 4E)}{2\sqrt{3}(1 + \cos 4E)} = \frac{D(E)}{\varrho},$$

Hamiltonian

$$H = 1/2 \left(p_\varrho^2 + \frac{(7 - \cos 4E)^2 p_E^2}{24\varrho^2(1 + \cos 4E)} \right) - \frac{1}{\varrho} \frac{\sqrt{7 - \cos 4E} (9 + \cos 4E)}{2\sqrt{3}(1 + \cos 4E)}$$

Parametrization

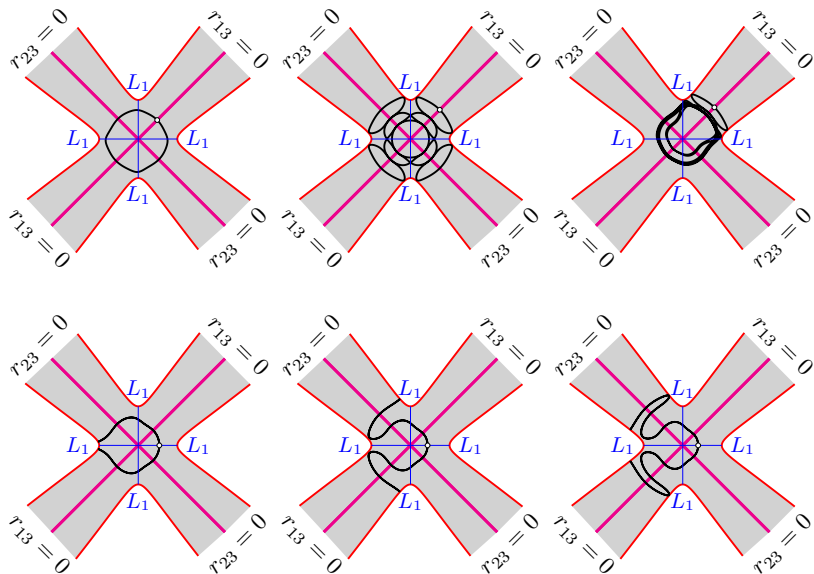
Not singular Hamiltonian

$$\begin{aligned} H' &= (1 + \cos 4E)(H - h) = \\ &= \frac{1}{2}(1 + \cos 4E)(p_\rho^2 - 2h) + \frac{(7 - \cos 4E)^2 p_E^2}{48\rho^2} - \frac{\sqrt{7 - \cos 4E}(9 + \cos 4E)}{2\sqrt{3}\rho}, \end{aligned}$$

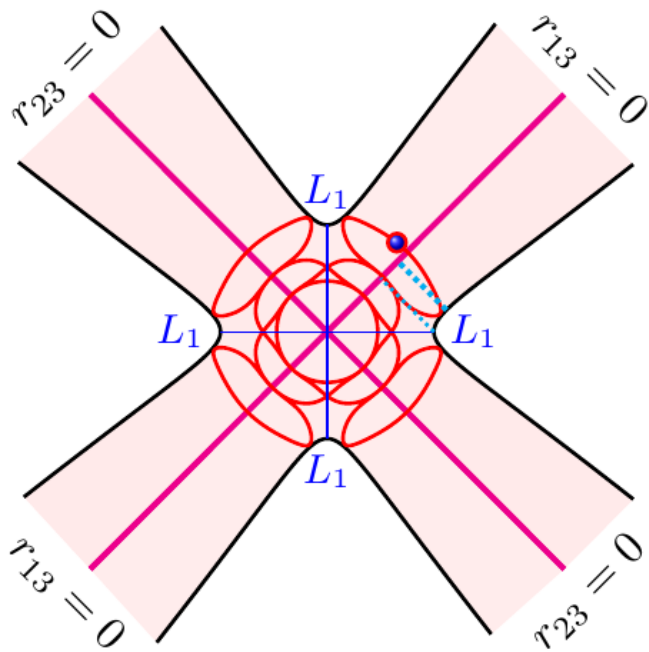
On the collision ray $E = \pi/4 + k\pi/2$:

$$p_E^2 = 2\sqrt{6}\rho.$$

Rectilinear trajectories, results



$$\rho = 3.74531687$$



Properties of rectilinear trajectories

- $\dot{E} = 0, E = 0 \Rightarrow$ homothetic motion
- All other trajectories intersect the rays $E = \pi/4 + k\pi/2$ и $E = k\pi/2$
- Free-fall trajectories are orthogonal to the zero velocity curve
- Rectilinear trajectories are orthogonal to collision rays
- The equality $p_E^2 = 2\sqrt{6}\rho$ holds on the collision rays $E = \pi/4 + k\pi/2$
- The maximum of ρ is achieved together with the maximum of p_E
- The intersection with Eulerian lines occurs between a series of collisions
- $\lim_{\rho \rightarrow \infty} l_\theta = \sqrt{2}; \lim_{\rho \rightarrow \infty} l_E \sim \sqrt{\rho}$.
- $\inf l_\theta$ is achieved when $\rho \rightarrow \infty$
 $\min l_E$ is achieved when $E = \pi/4 + k\pi/2 \pm 0.605632$ and is equal to $l_E \approx 2.285948$.

Chaos

- Each orbit intersects the ray $E = \pi/4$ and $p_{E_0}^2 = 2\sqrt{6}\rho_0$, the set of all orbits are defined by the number $\rho_0 \in \mathbb{R}^+$
- If initial point is located on the ray $E = 0$, then initial conditions are located in the region

$$\mathfrak{E} : 0 < \rho \leq \sqrt{5}/\sqrt{2}, p_{E_0}^2 \leq \frac{4\rho(5\sqrt{2} - 2\rho)}{3}.$$

The mapping $\mathfrak{T} : (0, \infty) \rightarrow \mathfrak{E}$ translates one-dimensional line to two-dimensional set \mathfrak{E} (similar to Peano curve?)

Isosceles trajectories

Parametrization:

$$z_E = \frac{\sqrt{2} \cos E}{\sqrt{3} - \cos E} + i \frac{\sqrt{3} \sin E}{\sqrt{3} - \cos E}.$$

Coordinates

$$\begin{aligned}\cos \theta &= \frac{(\cos 2E - 5)(1 + 3 \cos 2E)}{3 \cos^2 2E + 2 \cos 2E + 11}, \\ \sin \theta &= \frac{-8\sqrt{3} \sin E \sin 2E}{3 \cos^2 2E + 2 \cos 2E + 11},\end{aligned}$$

Hamiltonian

$$\begin{aligned}H = 1/2 \left(p_\varrho^2 + \frac{(3 \cos^2 2E + 2 \cos 2E + 11)^2 p_E^2}{24 \varrho^2 (1 - \cos 2E)(3 + \cos 2E)^2} \right) \\ - \frac{1}{\varrho} \frac{\sqrt{3 \cos^2 2E + 2 \cos 2E + 11} (7 - 3 \cos 2E)}{\sqrt{6}(1 - \cos 2E)(3 + \cos 2E)},\end{aligned}$$

Isosceles trajectories, results

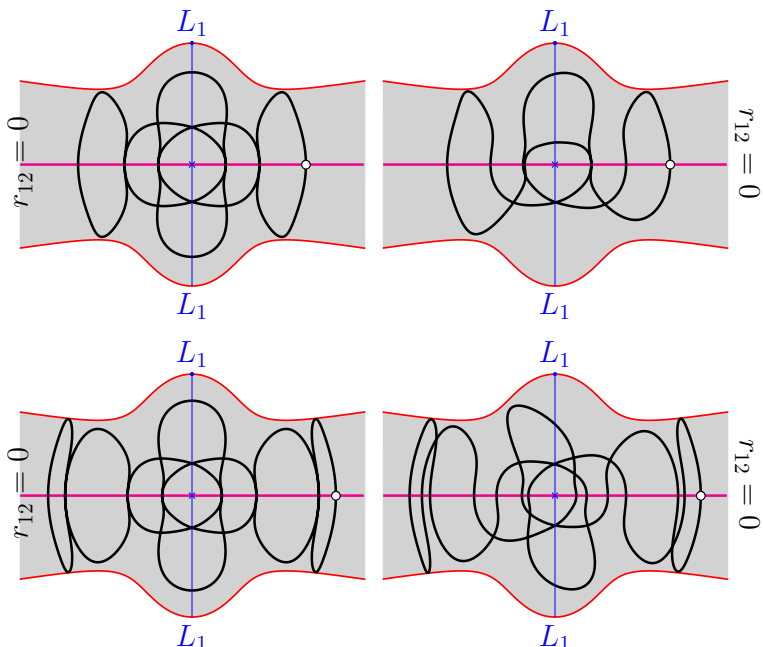
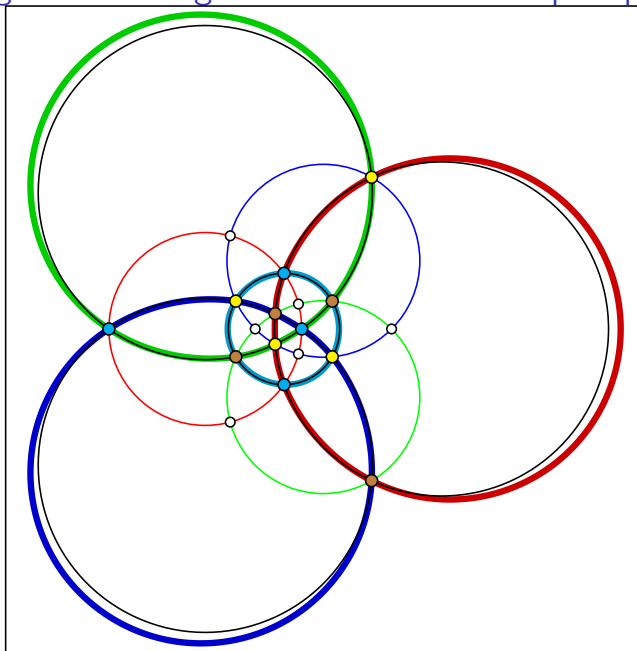
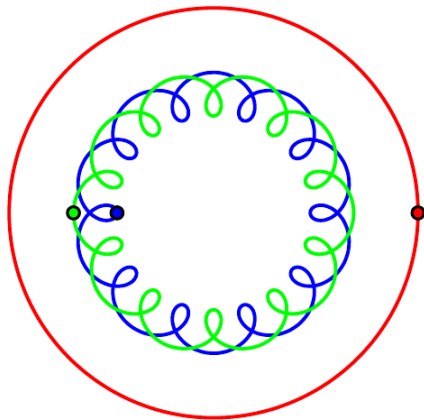
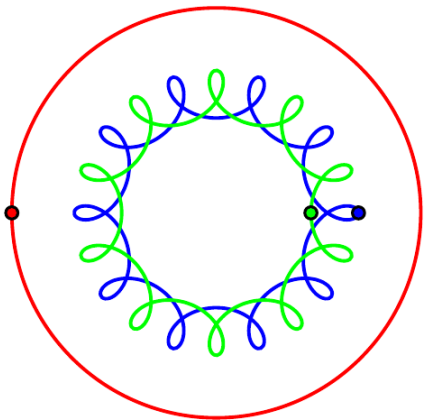


Figure-eight on the regularized extended complex plane





THANKS!!!